

PADRE MeDDEA

Steven Christe (NASA Goddard)

Measuring Directivity to Determine Electron Anisotropy (MeDDEA)

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SwRI Team: Amir Caspi, Carlos Urdiales, Reynaldo Gonzales

SO Team: Muriel Stiefel, Laura Hayes, Daniel Ryan, Natasha Jeffrey, Säm Krucker

and PADRE Team!



Fachhochschule
Nordwestschweiz



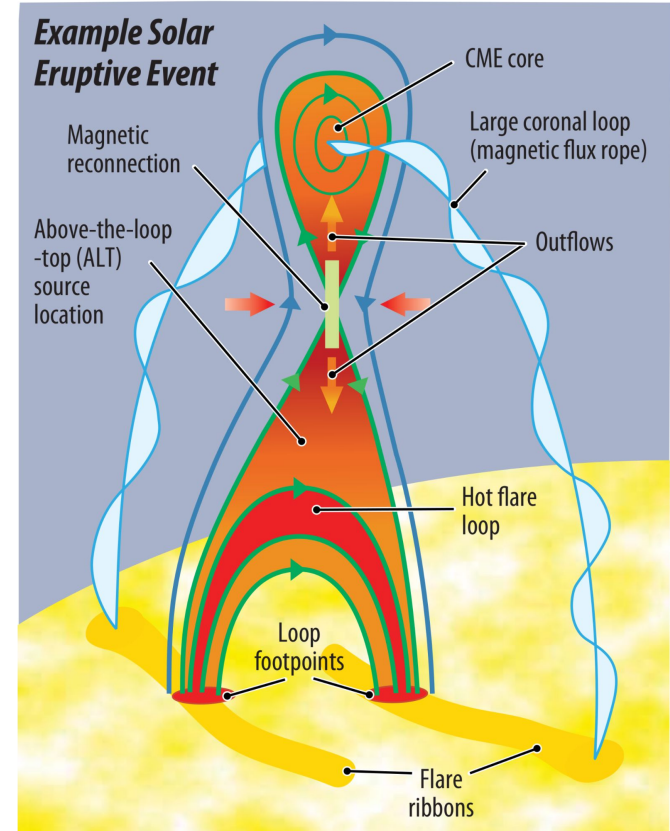
Particle Acceleration in Solar flares



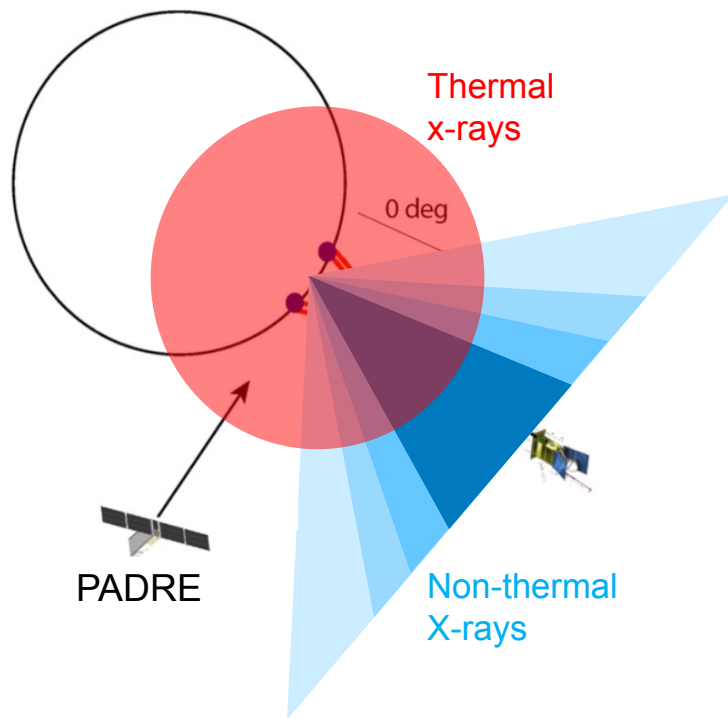
Solar flares are known to accelerate particle to high energies.

Some of these particles escape the Sun leading to intense and dangerous solar energetic particles (SEPs) that can damage satellites or endanger astronauts.

The physics of the acceleration is not yet understood limiting our ability to predict these dangerous events.



Measuring the X-ray directivity with PADRE and Solar Orbiter STIX



- MeDDEA makes energy-dependent measurements of X-ray flux from the Sun just like Solar Orbiter STIX
- Solar Orbiter provides opportunities to observe the same flares from a different angle than the Earth
- MeDDEA uses SO/STIX flight-spare detectors provided by CEA-Saclay.
- Comparing the two measurements, made using the same detectors in different locations will enable us to measure the X-ray directivity and constrain the acceleration mechanism of electrons.

What is PADRE?



PADRE is the Solar Polarization and Directivity X-Ray Experiment

12U CubeSat funded by the NASA

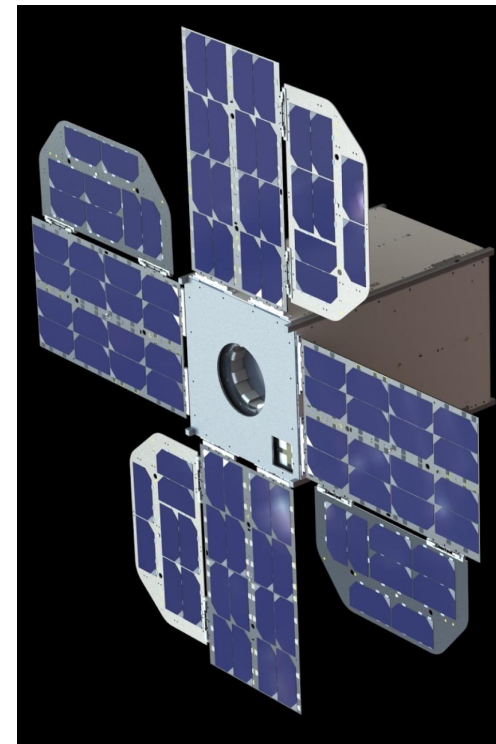
Led by UC Berkeley/Space Sciences Lab - Juan C. Martinez Oliveros (PI)

Science Goal: Investigate the dominant electron acceleration mechanisms in solar flares

Science Objective: Determine the degree of beaming (electron pitch angle anisotropy) of accelerated electrons in large solar flares

It achieves this single science objective with two instruments

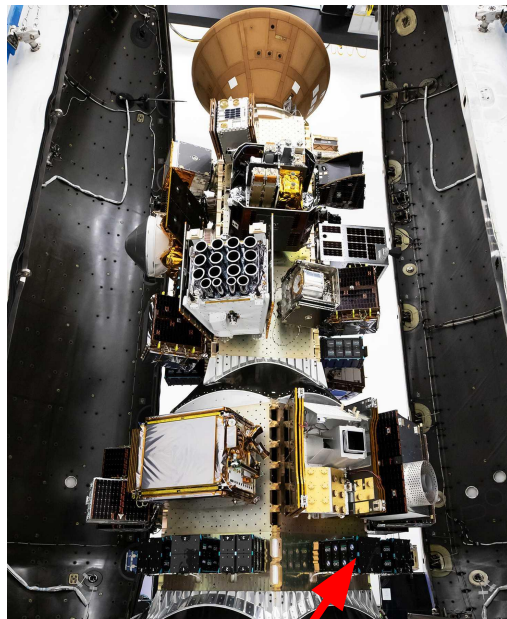
- **Measuring Directivity to Determine Electron Anisotropy (MeDDEA)**
- Solar Hard X-ray Polarimeter (SHARP)



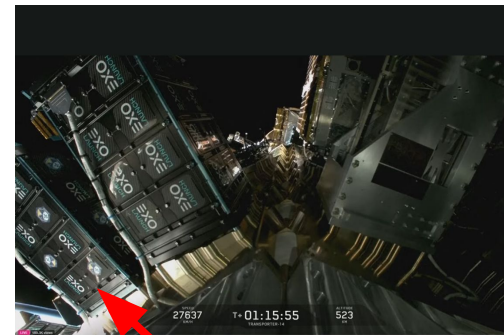
PADRE launched on 23-Jul-2025



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PADRE in
launcher



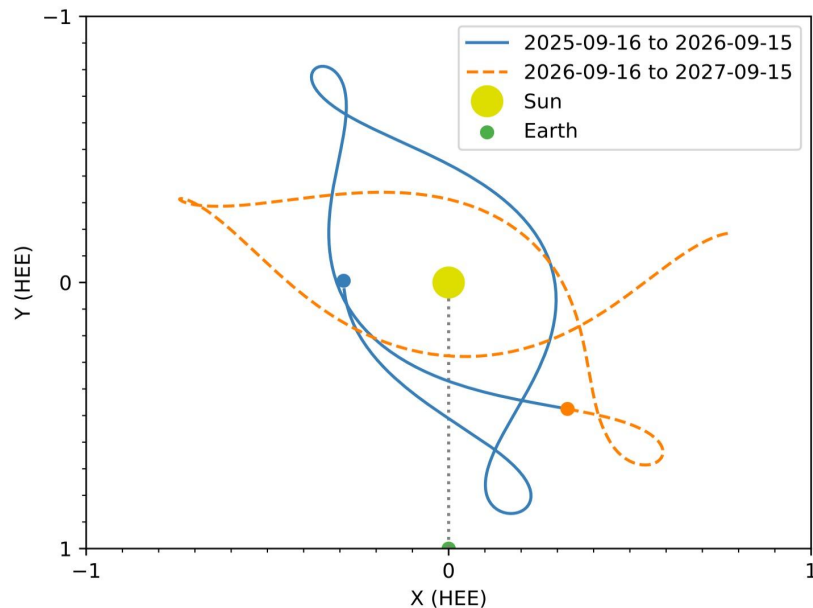
PADRE in
space before
deployment

NORAD ID: 64559
Perigee: 598.0 km
Apogee: 601.9 km
Inclination: 97.8 °
Period: 96.5 minutes
Semi major axis: 6970 km

Solar Orbiter STIX Orbit

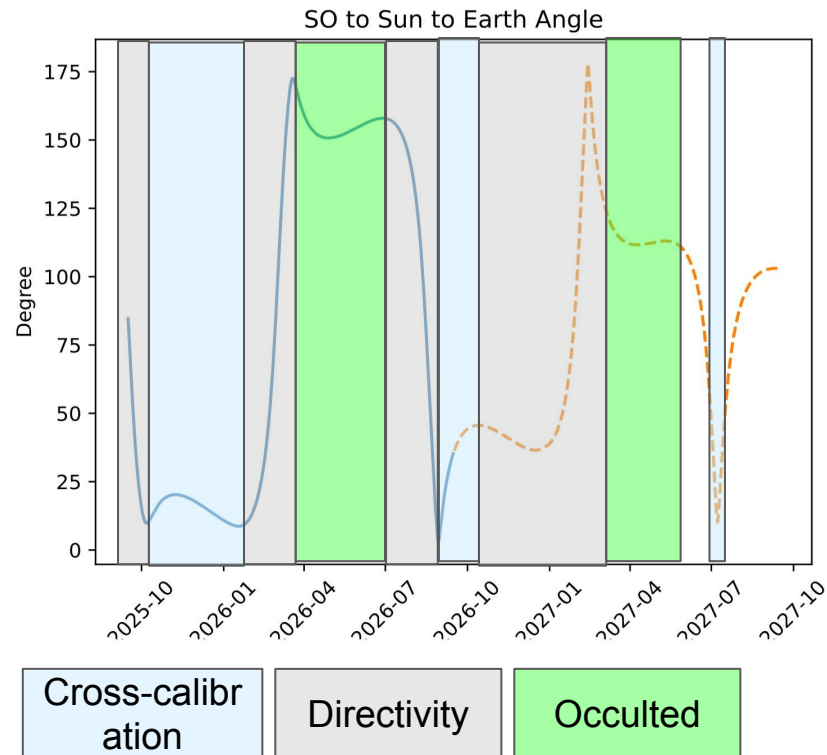


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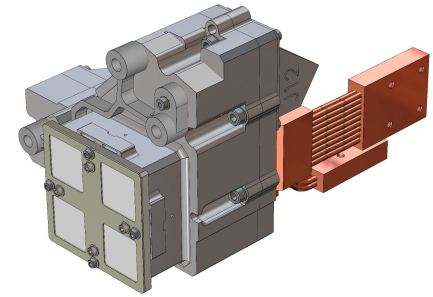
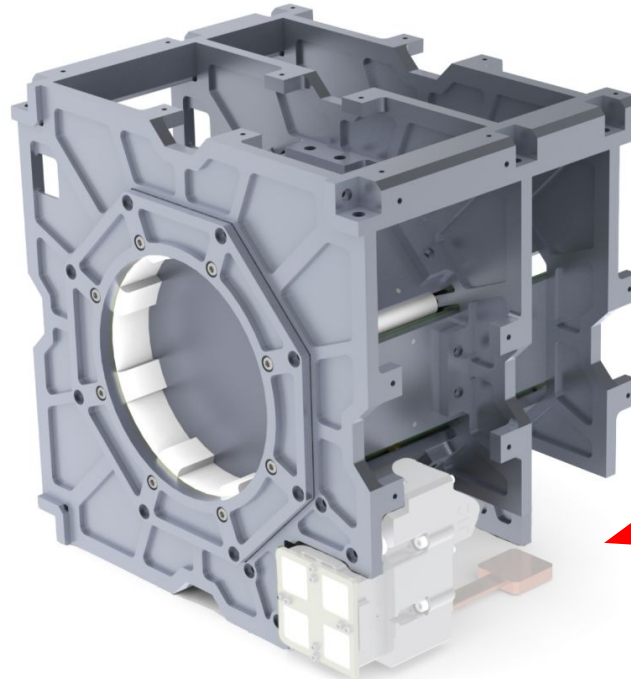
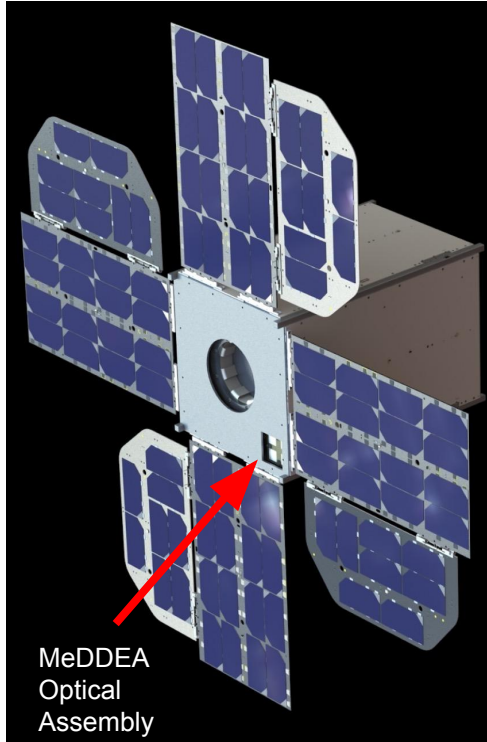


Blue - prime mission
orange - extended mission

2026 STIX Meeting



MeDDEA Optical Assembly on the Observatory

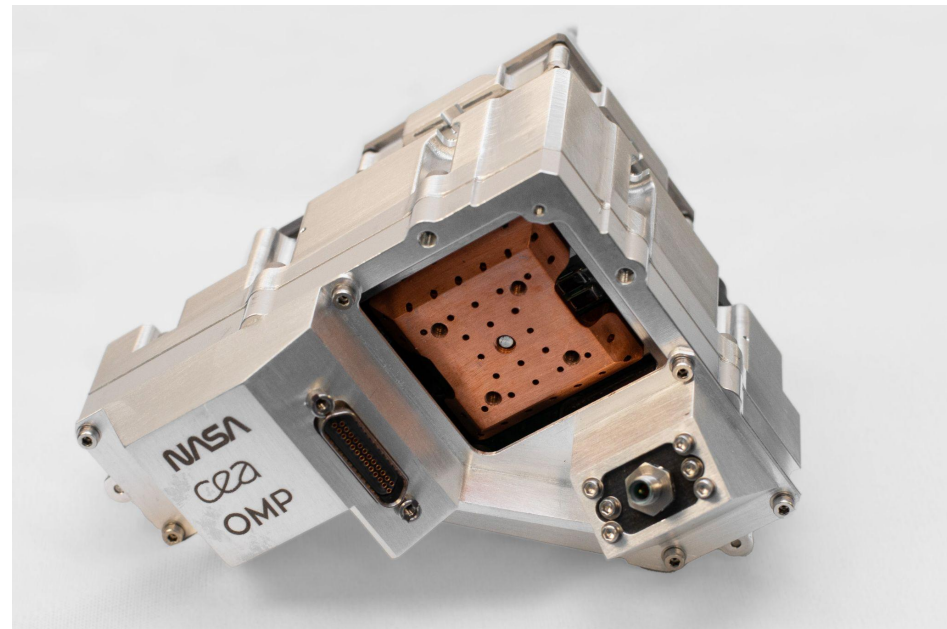
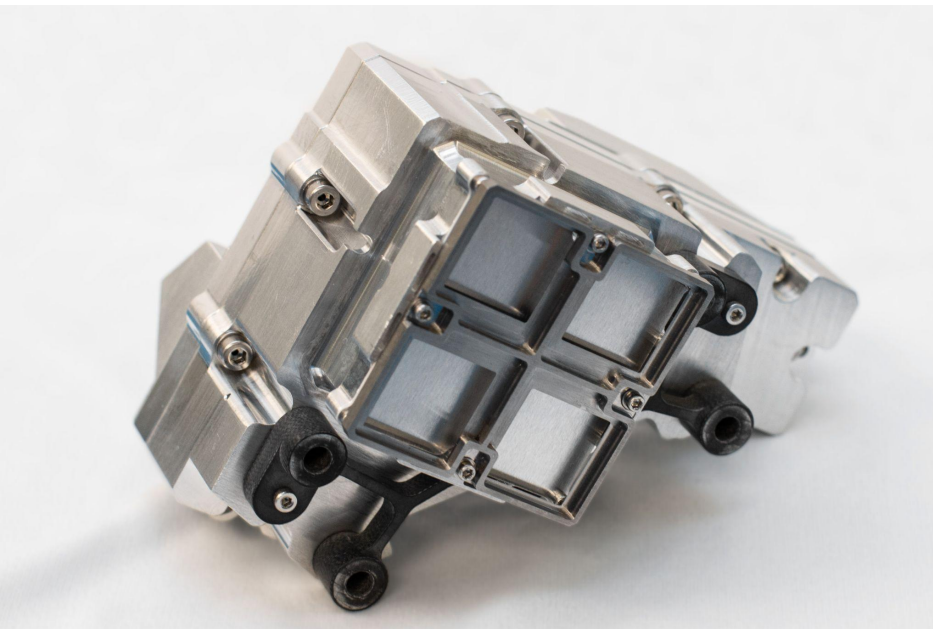


MeDDEA Optical Assembly

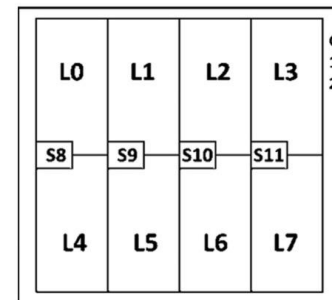
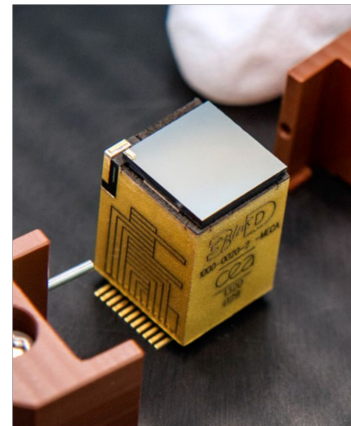
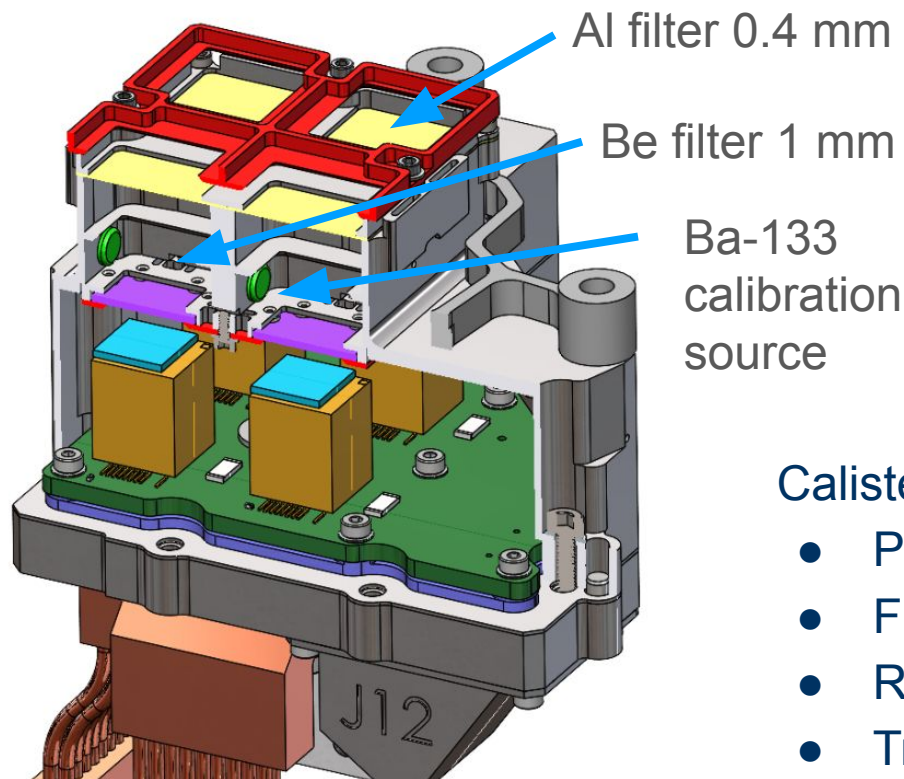
Optical Assembly Glamour Shots



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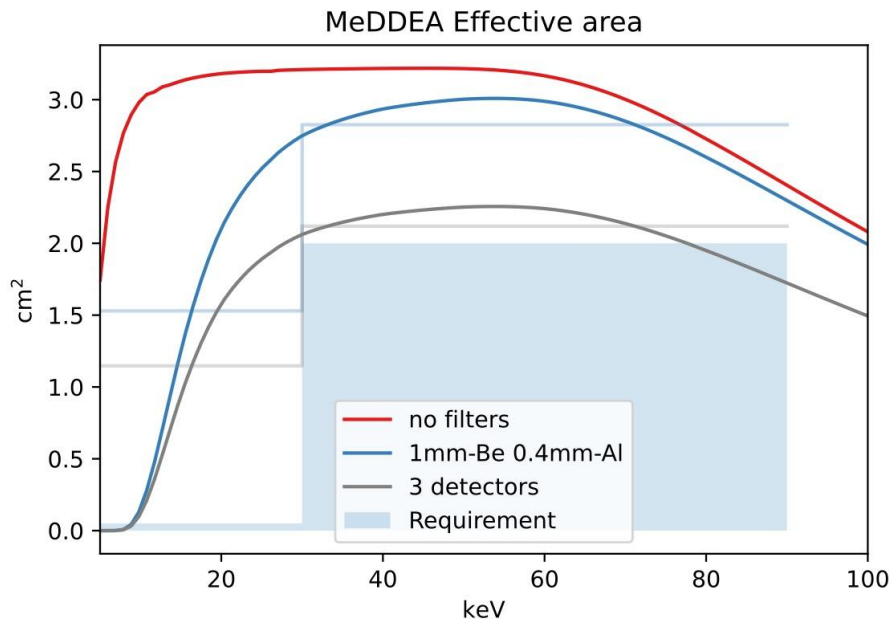
Optical Assembly



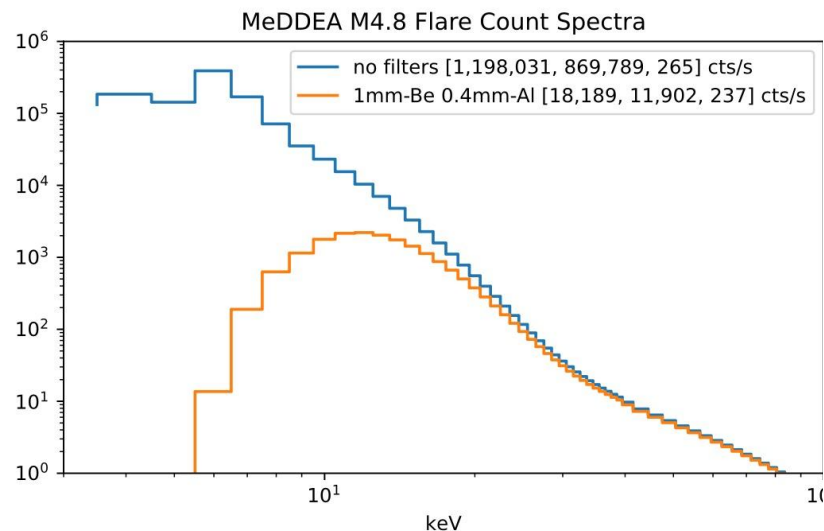
Caliste-SO Detectors

- Provided by CEA-Saclay
- Flight spares from Solar Orbiter STIX
- Rad-hard detectors and ASIC
- Triggered detector system, front-end custom ASIC (IDeF-X)

MeDDEA Effective Area



~ $\frac{1}{2}$ STIX effective area
32 detectors vs 4 detectors



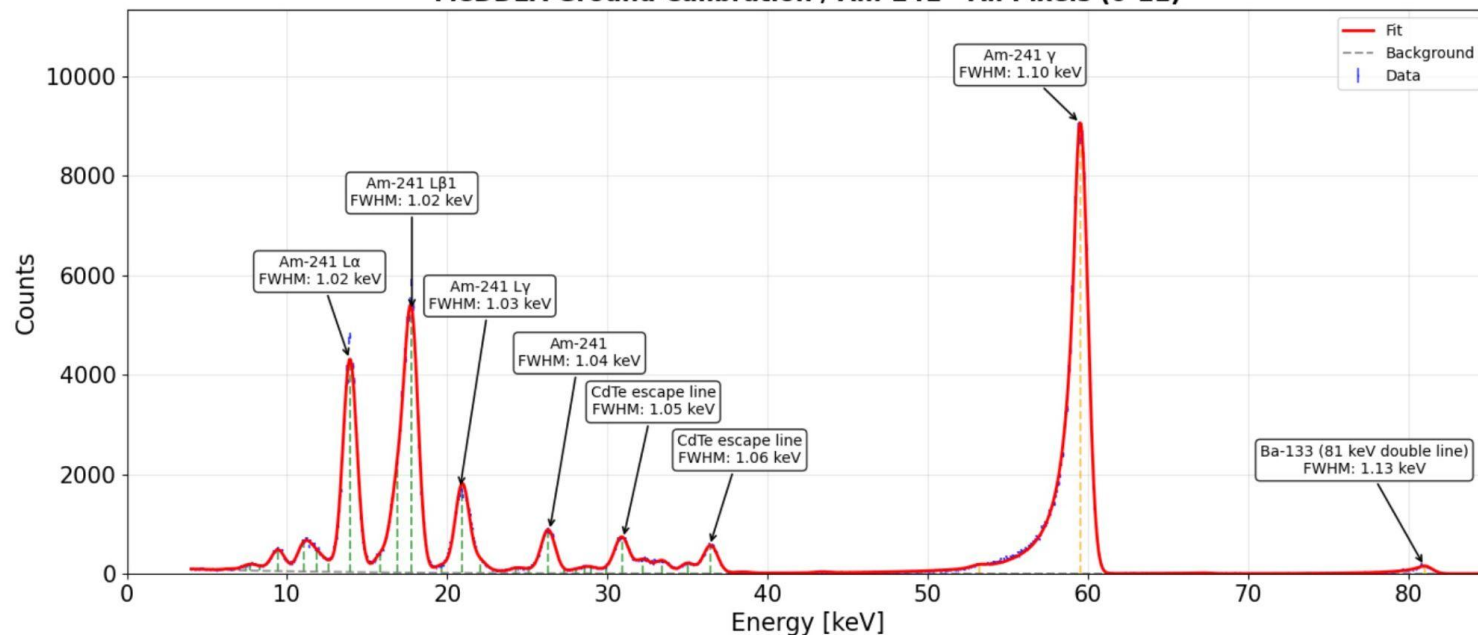
Why the thick filter?!

Ground Calibration



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MedDEA Ground Calibration / Am-241 - All Pixels (0-11)

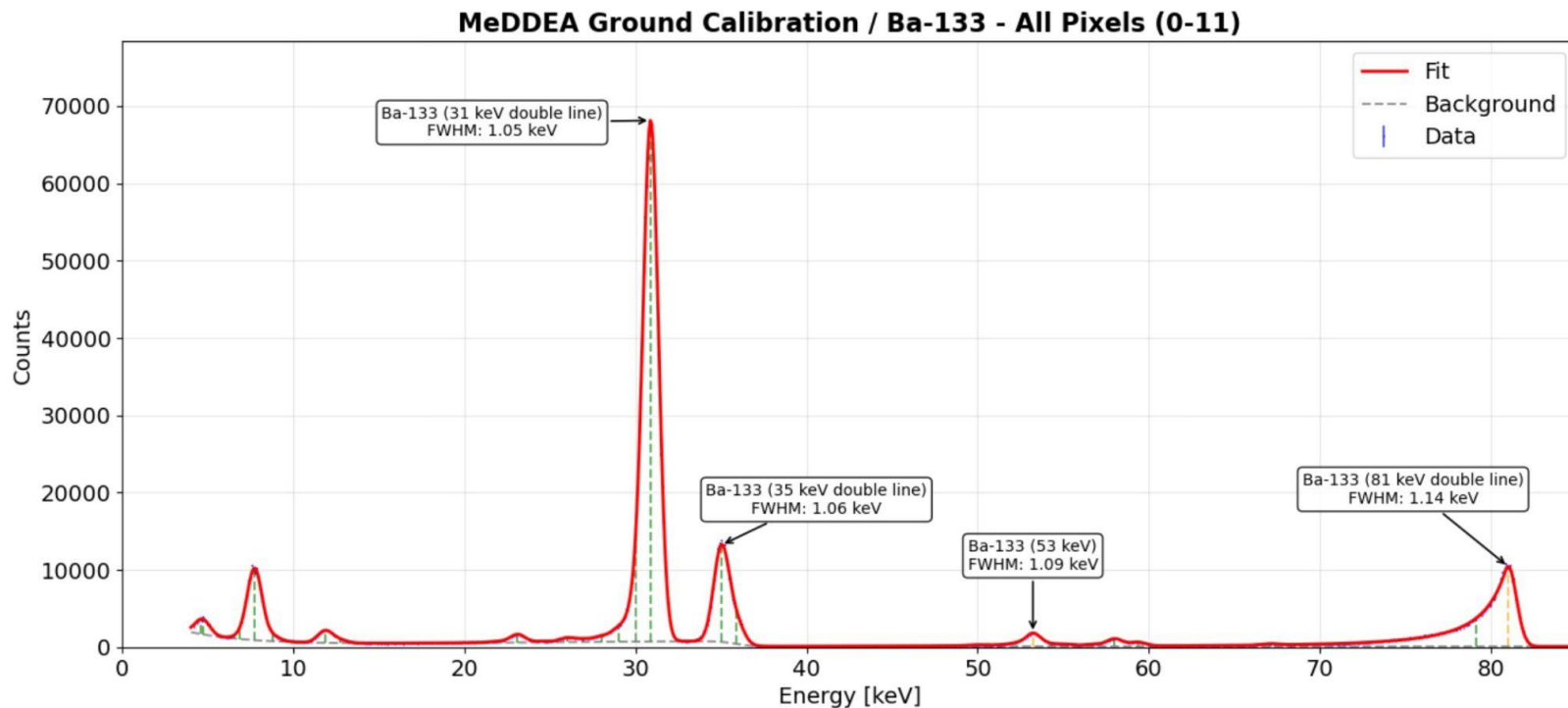


FWHM ~ 1 keV

Low energy
threshold ~ 3
keV

Flight
calibration is
comparable

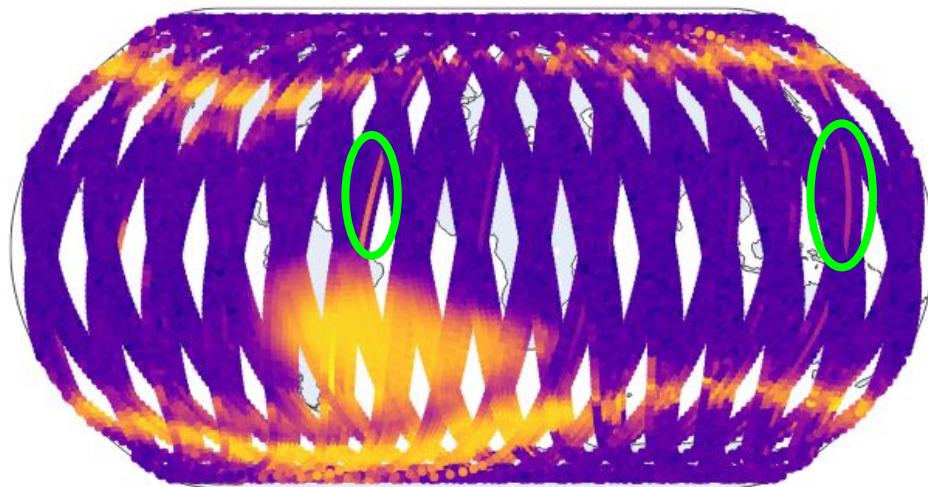
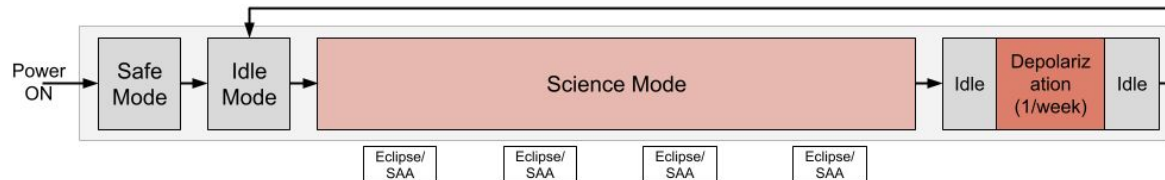
Calibration Spectrum



MeDDEA is “always” operating



MeDDEA Nominal Operations



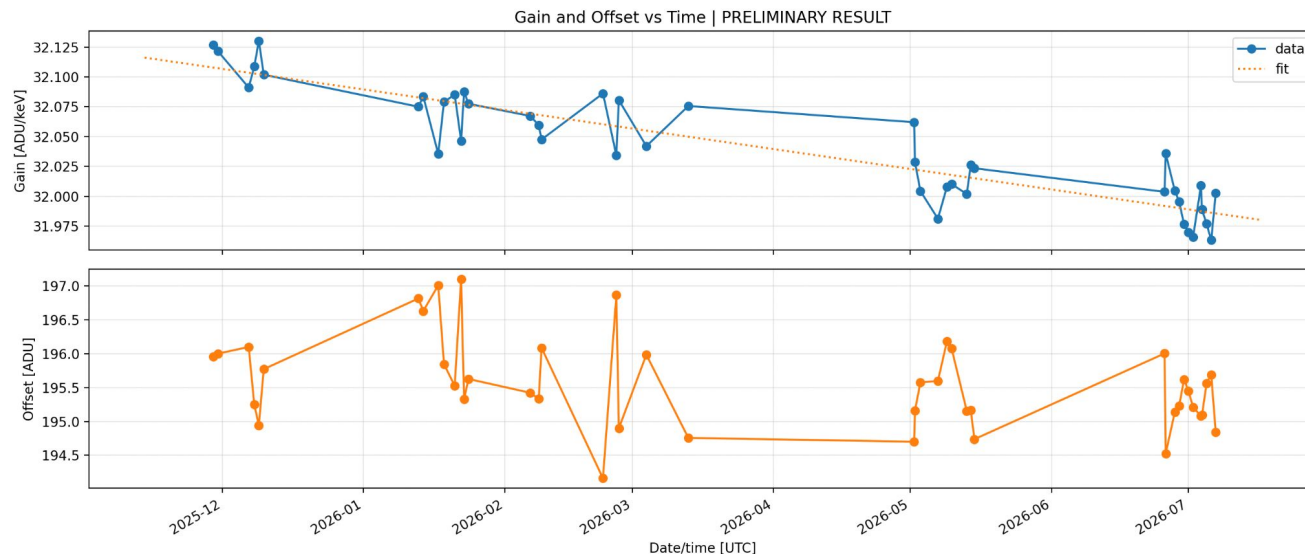
Take data throughout the orbit. In eclipse, measuring our calibration source.

Data also taken during SAA and belt crossing.

Potential monitor of space radiation in Low-Earth Orbit affecting spacecraft

Flares circled in green.

Performance Stability



- Preliminary look at detector gain/offset throughout the first year
- Very small changes seen and a slight trend over time.

Science Data Products



Detectors measure individual trigger events (photons) in each detector (time, pixel, energy). From this data, two products are generated.

- Raw event data (photon lists), stored and ~~downlinked on demand~~ **downloaded automatically**
 - Time, 12.8 us steps
 - All pixels
 - Energy bins ~ 0.02 keV steps
 - ~~Flare product~~
- Histogram/spectrum product (generated on board), ~~“always” available~~ **deprecated!**
 - Time, 10 s integration
 - 24/48 (selectable) pixels
 - Energy bins ~ 0.2 keV steps
 - Summary product (no decimation applied)

Changes in operations occurred $\sim 6/26/2026$

MeDDEA Detector Configurations



There are 3 detector configurations which control the detector thresholds

- **High Sensitivity**

- Lowest thresholds set. Primarily for calibration or quiet solar activity. All thresholds set to ~2.9 keV. **Primary mode**

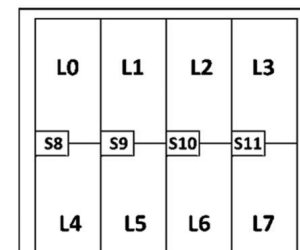
- **Default**

- 5 keV threshold for small pixels, 15 keV threshold for large pixels. **Primary mode** during active flaring times. Primarily for M class flares

- **Large Flares ($\geq$ X Class flares)**

- 5 keV threshold for small pixels. Most large pixels are turned off. Primary mode for extreme solar activity. **Never used.**

Changes in operations occurred ~6/26/2026



PADRE/MeDDEA Operations



56

Flares Observed

41 C-class • 15 M-class

466h

Science Data

MeDDEA downloaded

11

Months In Orbit

Launch: June 23, 2025

81

Days

SHARP powered on

91

Days

MeDDEA powered on

✓ **Minimum Mission Achieved** — Launched after maximum of SC25

- Decay Phase Ongoing • Big Flares Expected • First spinning sun-pointing CubeSat (1 rpm)

MeDDEA was powered on 24% of the time, and observed 14% of the time.

Observed ~41 C, ~15 M, 1X so far...

Operating a cubesat is difficult...we are improving!

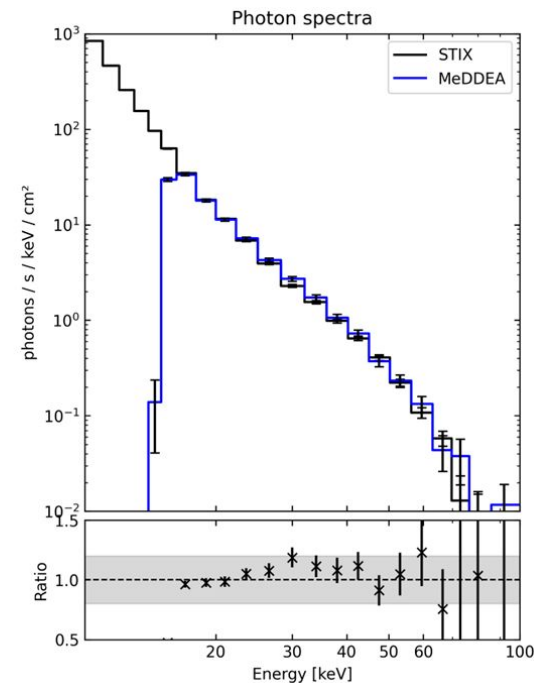
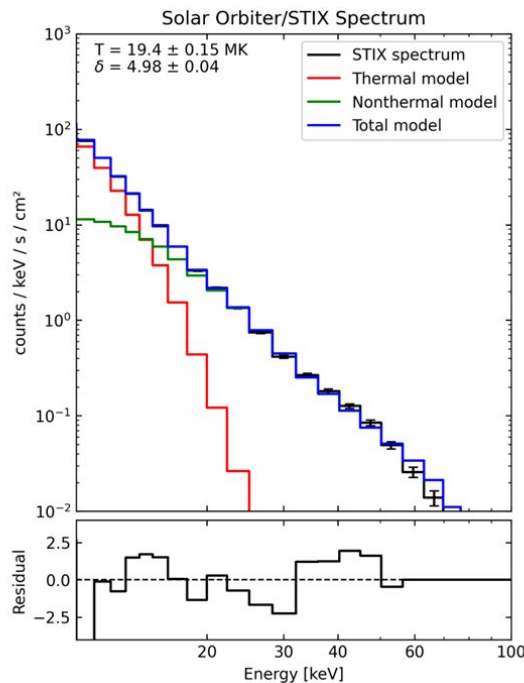
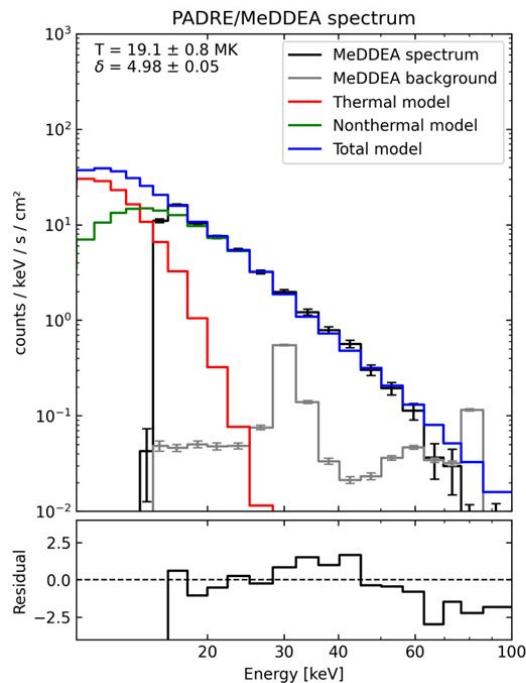
M1.6 2025-Nov-29 00:02:00



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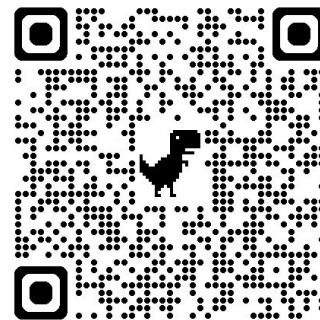
Cross-calibration between STIX and MeDDEA - 17-jan-2026 (M2 flare)



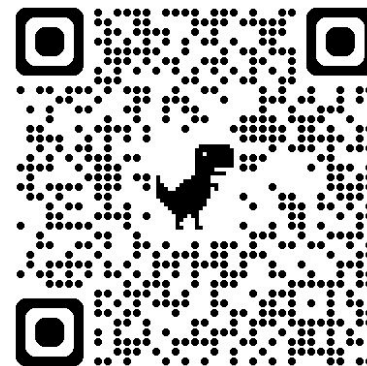
Work is ongoing but spectra can already be compared. **More about this in the next talk by M. Stiefel**

Data Availability

- Data is made publicly available as soon as it is downlinked.
- Data is publicly available now (raw, L0, L1) along with code to process the data.
Level 2 (calibrated data) to be released soon.
- MeDDEA data is available online from NASA SDAC:
<https://umbra.nascom.nasa.gov/padre/padre-meddea>
- Python package to analyze MeDDEA data
https://github.com/PADRESat/padre_meddea
- [MeDDEA Flare list](#)
- Contact me steven.christe@nasa.gov



Data Dashboard



GitHub Repository